currents of the fea have fooped out the valleys, elevated the hills, and beflowed on them their corresponding directions; that the fame waters of the ocean, by transfooring and depositing earth, See, have given rife to the parallel flrans; that the waters from the heavens gradually defretoy the effects of the fea, by continually diminishing the height of the mountains, filling up the valleys, and chooking the mouths of rivers; and, by reducing every thing to its formerievel, they will, in time, reflore the earth to the fea, which, by its natural operations, will again create now continents, interferefed with mountains and valleys, every way finalir to those which we now inhabit.

PROOFS

.

THEORY OF THE EARTH.

ARTICLE I.

Of the Formation of Planets.

A snatural history is our proper fubject, we call observations. But, as the earth is fo nearly related to the heavenly bodies, and, as observations of this kind illustrate more fully those doctrines we have already advanced, it is necessity to give form general ideas concerning the formation, motion, and figure, of the earth, and

other planets.

The earth is a globe of about 3000 leagues in diameter; it is fituated 30 million of leagues from the fun, round which it revolves in 365 days. This annual revolution is the effect of two forces; the one may be confidered as an impulse from right to left, or from left to right;

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the other as an attraction from above downwards, or from below upwards, to a common centre. The direction and quantity of these forces are combined, and fo nicely adjusted, that they produce a uniform motion in an ellipse approaching to a circle. Like the other planets, the earth is opaque, throws out a fhadow, and reflects the rays of the fun, about which it revolves in a time proportioned to its relative diftance and denfity. It likewise revolves about its own axis in 24 hours; and its axis is inclined to the plane of its orbit 66; degrees. Its figure is that of a fpheroid, the two axes of which differ from each other about an 165th part; and it revolves round the fhortest axis,

These are the principal phanomena of the earth, the refults of discoveries made by means of geometry, astronomy, and navigation. It is unnecessary here to enumerate the proofs and observations by which these facts have been eftablished. We shall confine our remarks to such objects as are still doubtful; and shall therefore proceed to give our ideas concerning the formation of planets, and the changes they have undergone, previous to their arriving at the state in which we now perceive them. To the many fystems and hypotheses which have been framed concerning the formation of the earth, and the different flates it has paffed through, we may be allowed to add our own conjectures, especially as we are determined to support them with a fuperior degree of probability; and we are the more encouraged to deliver our notions on this fubject, because we hope to enable the reader to diftinguish between an hypothesis which consists only of poffibilities, and a theory supported by facts; between a fystem, such as we are about to give, of the formation and primitive state of the earth, and a physical history of its real condition, which has been delivered in the prece-

ding difcourfe.

Galileo traced the laws of falling bodies; and Kepler observed, that the areas which the principal planets describe in moving round the fun, and those of the fatellites round their principal planets, were proportioned to the periods of their revolutions, and that these periods were as the fquare roots of the cubes of their diffances from the fun, or from the principal planets. Newton discovered that the power of gravity extended to the moon, and retained it in its orbit : that the force of gravity diminished in exact proportion to the fquares of the diffances, and, confequently, that the moon is attracted by the earth : that the earth, and all planets, are attracted by the fun; and, in general, that all bodies which revolve about a centre, and deferibe areas proportioned to the periods of their revolution, are attracted by that luminary. Gravity, therefore, is a general law of nature. The planets, comets, the fun, the earth, are all fubject to its laws; and it is the fource of that harmony mony which prevails in the univerfe. Nothing in physics is better elabilished than the existence of this power in every material body. Repeated experience has consirmed the effects of its insistence, and the labour and ingenuity of geometers have determined its quantity and relations.

This general law being once discovered, the effects of it would be eafily explained, if the action of those bodies which produce them were not too complicated. A flight view of the folar fystem will convince us of the difficulties which attend this fubiect. The principal planets are attracted by the fun, the fun by the planets, the fatellites by the principal planets, and each planet attracts all the others, and is attracted by them. All these actions and re-actions vary according to the quantities of matter and the diftances, and give rife to great inequalities, and even irregularities. How are fo many relations to be combined and estimated? Among such a number of objects, how is it possible to trace any individual? These difficulties, however, have been furmounted: the reasonings of theory have been confirmed by calculation; every observation has produced a new demonstration; and the fyftematic order of the universe is now laid open to every man who is able to diffinguish truth from error.

The force of impulsion, or what is commonly called the centrifugal force, is still unknown; but it affects not the general theory. It is evident, that, as the attractive force continually draws all the planets towards the fun, they would fall in a perpendicular line into that luminary, if they were not kept at a distance by fome other power, forcing them to move in a ftraight line. If, again, this impulfive force were not counteracted by that of attraction, all the planets would fly off in the tangents of their respective orbits. This progressive or impulsive force was unquestionably at first communicated to the planets by the Supreme Being. But, in phylical fubjects, we ought, as much as possible, to avoid having recourse to supernatural causes ; and, I imagine, a probable reason may be asfigned for the impulfive force of the planets, which will be agreeable to the laws of mechanics, and not more furprifing than many revolutions that must have happened in the universe.

The figher of the fun's attraction is not limited by the orbits of the planets, but extends to an indefinite diffance, always decreafing according as the figures of the augmented diffances. The comets, it is evident, which fedape our fight in the heavenly regions, are, like the planets, fully of the first of the fir

extent.

extent. The motions, therefore, both of planets and comets, are regulated by impulsive and attractive forces continually acting upon them, and obliging them to describe curves. But it is worthy of remark, that comets run through the fystem in all directions; that the inclinations of the planes of their orbits are fo very different, that though, like the planets, they be subject to the law of attraction, they have nothing in common with regard to their progressive or impulfive motions, but appear, in this respect, to be abfolutely independent of each other. The planets, on the contrary, move round the fun in the fame direction, and nearly in the fame plane, the greatest inclination of their planes not exceeding 7; degrees. This fimilarity in the polition and motion of the planets indicates, that their impulfive and centrifugal forces must have originated from one common cause.

May we not conjecture, that a comet falling into the body of the fun might drive off fome parts from its furface, and communicate to them a violent impultive force, which they full retain? This conjecture appears to be as well founded as that of Mr Leibnitz, which fluppoded the earth and planets to have formerly been funs; and his fyftem, of which an abridgment will be given in Art. V. would have been more comprehenfive, and more confonant to probability, if it had embraced the above idea. We agree with him, that this effect was produced at

the time when God is faid by Mofes to have feparated the light from the darknefs; for, according to Leibnitz, the light was feparated from the darknefs when the planets were extinguished. Bur, on our fupposition, there was a real physical feparation; because the opaque bodies of the planets were detached from the luminous matter of which the fain is composited.*

This notion concerning the cautie of the centrifugal force of the planets will appear to be lefs exceptionable, after we have collected the analogics, and effinanted the degrees of probability by which it may be fupported. We final first mention, that the motion of the planets have one common diredion, namely, from west to easily the case of the case o

This probability will be greatly augmented, if the confider the finalizary in the inclinations of the planes of their orbits, which exceed not 7½ degrees; for, by calculations, it has been differenced, that it is 24 to 1 againft any two planets being found, at the fame time, in the most being found, at the fame time, in the most plane, and their orbits; and, confequently, 24%, no 76926744 to 1, that this effect could not

^{*} If the Count de Buston had known, that the nucleus of the fun was a folid and opaque matter, a diffeovery lately made by the ingestious Dr. Willon of Glafgow, his hypothefis would have laboured under fewer difficulties.

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be produced by accident; or, what amounts to the fame, there is this great degree of probability, that the planets have been impressed with one common moving force, from which they have derived this fingular position. But nothing could bestow this common centrifugal motion, except the force and direction of the bodies by which it was originally communicated. We may, therefore, conclude, that all the planets have probably received their centrifugal motion by one fingle stroke. Having established this degree of probability, which almost amounts to a certainty, I next inquire what moving bodies could produce this effect; and I can find nothing but comets capable of communicating motion to fuch vaft maffes.

Upon examining the course of comets, it is eafy to believe that fome of them must occafignally fall into the fun. The comet 1680 approached fo near, that, at its perihelion, it was not more distant from the sun than a fixteenth part of his diameter; and, if it returns, which is extremely probable, in the year 2255, it may then fall into the fun. This event must depend upon the accidents it meets with in its courfe, and the retardations it fuffers in paffing through the fun's atmosphere *.

We may, therefore, prefume, with the great Newton, that comets fometimes fall into the fun. But they may fall in different directions.

If they fall perpendicularly, or in a direction not very oblique, they will remain in the body of the fun, ferve the purposes of fewel, and, by their impulse, remove the fun from his place, in proportion to the quantity of matter they contain*. But, if a comet falls in a very oblique direction, which will most frequently happen, it will only graze the furface, or penetrate to no great depth. In this cafe, it may force its way past the fun, detach certain portions of his body, to which it will communicate a common impulsive motion; and these portions pushed off from the fun, and even the comet itself, may turn planets, which will revolve round this luminary in the fame direction, and nearly in the fame plane. A calculation, perhaps, might be made of the quantity of matter. velocity, and direction, a comet ought to have, in order to force from the fun maffes equal to those which compose the fix planets and their fatellites. But it is fufficient here to observe. that the whole planets, with their fatellites. make not a 650th part of the fun's maist; for although the denfity of Saturn and Jupiter be less than that of the fun, and though the earth be four times, and the moon near five times, more denfe than the fun; yet they are only atoms when compared to his immenfe volume.

^{*} Oner. Would not fuch an event, by assementing the fun's quantity of matter, and, confequently, his attractive power, produce other changes in the folar fyflem? + See Newt. p. 405. E 2

^{*} See Newt. edit. 3. p. 525.

It must be acknowledged, that, although a 650th part of a whole may feem inconfiderable, it would require a very large comet to detach this part from the fun. But, if we confider the prodigious rapidity of comets in their perihelion; the near approach they make to the fun; the denfity and firong cohelion of parts neceffary to fullain, without destruction, the inconceivable heat they undergo; and the folid and brilliant nucleus which shines through their dark atmospheres; it cannot be doubted that comets are composed of matters extremely dense and folid; that they contain, in fmall limits, a great quantity of matter; and, confequently, that a comet of no enormous fize may remove the fun from his place, and give a projectile motion to a mass of matter equal to the 650th part of his body. This remark corresponds with what we know concerning the respective densities of the planets, which always decrease in proportion to their distances from the fun, having less force of heat to refift. Accordingly, Saturn is less dense than Jupiter, and Jupiter much less than the earth. Thus, if the denfity of the planets, as Newton alleges, be in proportion to the quantity of heat they fupport, Mercury will be feven times denfer than the earth, and 28 times denfer than the fun, and the comet 1680, 28,000 times more denfe than the earth, or 112,000 times denfer than the fun. Now, fuppoling the quantity of matter in this comet to be equal to a

ninth part of the fun, or allowing it to be only roodth part of the bulk of the earth, its quantity of matter would still be equal to a gooth part of the fun: Hence a body of this kind, which would be but a fmall comet, might push off from the fun a gooth or a 650th part, especially when the amazing rapidity of comets, in their perihelion, is taken into the calculation.

The correspondence between the density of the whole planets, and that of the fun, deferves alfo to be noticed. Upon, and near the furface of the earth, there are fubftances 1400 or 1500 times denfer than others; the denfities of air and gold are nearly in this proportion. But the interior parts of the earth and planets are more uniform, and differ little with regard to denfity; and the correspondence in the density of the planets and that of the fun is fo great, that, out of 650 parts, which comprehend the whole denfity of the planets, there are more than 640 nearly of the fame denfity with the folar matter: and there are only ten of those 650 which are of a fuperior denfity; for the denfity of Saturn and Jupiter is nearly the fame with that of the fun; and the quantity of matter in those two planets is at least 64 times greater than what is contained in the four inferior planets, Mars, the Earth, Venus, and Mercury. We may, therefore, conclude, that, in general, the matter of the planets is very nearly of the fame

kind with the folar matter, and, of course, that the former may have been separated from the

To this theory it may be objected, that, if the planets had been driven off from the fun by a comet, in place of deferibing circles round him, they must, according to the law of projectiles, have returned to the fame place from whence they had been forced? and, therefore, that the projectile force of the planets cannot be attributed to the impulse of a comet.

I reply, that the planets issued not from the fun in the form of globes, but in the form of torrents, the motion of whose anterior particles must have been accelerated by those behind, and the attraction of the anterior particles would alfo accelerate the motion of the posterior; and that this acceleration, produced by one or both of these causes, might be such as would necessarily change the original motion arising from the impulse of the comet, and that, from this cause, might refult a motion fimilar to what takes place in the planets, especially when it is considered, that the shock of the comet removes the sun out of its former flation. This reasoning may be illustrated by an example. Suppose a musketball discharged from the top of a mountain, and that the force of the powder was fufficient to push it beyond a semidiameter of the earth, it is certain that this ball would revolve round

the earth, and return at every revolution to the place from whence it had been discharged. But, inflead of a musket-ball, if a rocket were employed, the continued action of the fire would greatly accelerate the original impulsive motion. This rocket would by no means return to the fame point, like the ball; but, cæteris paribus, would describe an orbit, the perigee of which would be more or less distant from the earth in proportion to the greatness of the change produced in its direction by the accelerating force of the fire. In the fame manner, if the original projectile force impressed by the comet on the torrent of folar matter was accelerated, it is probable, that the planets formed by this torrent acquired their prefent circular or elliptical movements around the fun-

The appearance exhibited in great cruptions from volcano's may give forne idea of this acceleration of motion. When Vetivius begins to groun and throw out inflamed matter, it has been often remarked, that the motion of the cloud first ejecked is slower than the succeeding ones, and that they go on interesting in celerity, till at last sliphur, lava, melted metal, and huge stones are thrown up; and that, though these observe nearly the same direction, they alter considerably that of the first cloud, and clevate it to a greater height than it would otherwise have reached.

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The objection will be fill further obvirate, if its condidered, that the inpulle of the comet into continue, that the inpulle of the comet mul; in some degree, have communicated a motion to the finn, and removed it from its former fination; and that, although this motion may now be fin final as to efcape the notice of affections, it may fill, however, exitl, and make the finn decirbs a curve round the centre of gravity of the fyftem. If this be allowed, as I prefine it will, the planes, instead of returning to the finn's body, would deferibe orbits, the perihelions of which would be as diltant from the finn as the space that he presently occupies is distant from the soriginal flattion.

It may be further objected, that, if motion be accelerated in the fume direction, no change in the perhicition could take place. But is it credible, that no change of direction can happen in a torrent whose partices fuceed each other? On the contrary, it is extremely probable, that a change was actually produced infi-ficient to make the planets move in their present

It may fill be objected, that, if the fituation of the fun had been changed by the shock of a comet, it would move uniformly; and, of course, this motion being common to the whole system, no alteration would be effected. But, previous to the shock, might not the sun move round the centre of the cometary system; and might

not

not this primary motion be augmented or diminished by the stroke of the comet? Is not this sufficient to account for the actual motion of the planets?

If none of these suppositions be admitted, may it not be prefumed, that the elafticity of the fun might elevate the torrent above his furface, in place of pushing it directly forward? This circumfrance alone would be fufficient to remove the perihelion, and endow the planets with their present movements. Neither is this supposition destitute of foundation: The folar matter may be exceedingly elaftic; fince light, the only part of it we are acquainted with, feems, by its effects, to be perfectly elaftic. I acknowledge that I cannot determine which of the causes above affigned has actually produced an alteration in the projectile force of the planets; but they at least show that such a change is not only possible, but probable, which is enough for my present purpose.

Without fartiler infifting on the objections which may be made againft my hypothetis, or the analogical proofs that might be brought in fupport of it, I final profecute my fubject, and draw the proper conclutions. Let us first examine what might happen to the planets, and particularly to the earth, when they were impressed with their projectile forces, and what was their flate after their feparation from the body of the

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fun. A projectile motion having been communicated, by the firoke of a comet, to a quantity of matter equal to a 650th part of the fun's mass, the light particles would separate from the denfe, and, by their mutual attractions, form globes of different folidities. Saturn being composed of the largest and lightest parts, would be removed to the greatest distance from the fun; Jupiter, being denfer than Saturn, would have a nearer station; and so of the rest. The largest and least folid planets are most distant, because they received a greater projectile force than the fmaller and denfer; for the projectile force being proportioned to the furface to which it is applied, the fame stroke would make the larger and lighter parts of the folar matter move with more rapidity than the fmaller and heavier. The parts, therefore, which differed in denfity, would feparate from each other in fuch a manner, that, if the denfity of the folar matter be equal to 100, that of Saturn will be equal to 67, of Jupiter, = 941, of Mars, = 200, of the Earth, = 400, of Venus, = 800, of Mercury, = 2800. But, as the attractive force acts not in proportion to the furface, but to the quantity of matter, it would retard the progress of the more dense parts of the folar matter; and it is for this reason that we find the most dense planets nearest the fun, and which move round him with more rapidity than those that are more diftant, and less dense.

The denfity and projectile motion of Saturn and Juniter, the two largest planets in the system. have the most exact proportion. The density of Saturn is to that of Jupiter as 67 to 041, and their velocities are nearly as 884 to 1204, or as 67 to 0011. How rarely do pure conjectures correspond to exactly to the phænomena of nature? It is true, according to this relation between the celerity and denfity of the planets. the denfity of the earth ought not to exceed 206,2, inflead of 400, which is its real denfity: hence it may be supposed, that the earth has now double its original denfity. With regard to the other planets, Mars, Venus, and Mercury, as their denfities are only conjectural, we know not whether this circumstance would confirm or weaken our hypothesis. Newton says, that the denfities of the planets are proportioned to the degrees of heat they are exposed to; and it is in confequence of this idea that we have mentioned Mars as being one time lefs denfe than the earth, Venus one time, Mercury feven times, and the comet 1680, 28,000 times denfer than the earth. But, if we attend to Saturn and Jupiter, the two principal planets, we will find, that this supposed proportion between the denfities of the planets, and the heat they fuftain, is not well founded: For, according to this hypothefis, the denfity of Saturn would be as 4.4. and that of Jupiter as 1417, instead of the proportions of 67 and 041; differences fo great as to delivoy the principles upon which they are founded. Thus, notwithflanding the regard due to the conjectures of Newton, I canner refrini from thinking that the denfities of the planets have a nearer relation to their celerities than rio the degrees of heat to which they are exported. This, indeed, is only a final caule; but the other is a phyrical relation, the exactness of which is remarkable in Saturn and Ipplier. It is certain, however, that the denfity of the earth, inflead of being 20%, is 400; and, confequently, the earth much have fuffered a condensation in the proportion of 200 to 400.

But have the condensations of the planets no relation to the quantity of folar heat they fuftain? In that cafe, Saturn, which is at the greatest distance from the sun, would have suffered little or no condensation; and Jupiter would be condenfed from 9011 to 941. Now, the fun's heat in Jupiter being to his heat in the earth as 1417 to 400, their condensation ought to be in the same proportion. Thus, if Jupiter be condensed as 9011 to 941, the earth, if it had been in the orbit of Jupiter, would have been condensed from 2067 to 215 990; but the earth being much nearer the fun, and receiving heat, in proportion to that of Jupiter, as 400 to 14:7, the quantity of condensation it would have undergone in the orbit of Jupiter must be multiplied by the proportion of 400 to 14:7, which will give nearly 234; for the condenfation the earth must have received. The denfity of the earth was coft; by adding its acquired condentation, its actual density will be 4001, which is nearly the same with 400, its real density determined by the moor's parallax. With regard to the other planets, I pretend not to give exact proportions, but only approximations, tending to shew, that their densities have a strong connection with the celerity of their motions in their respective orbits.

The comet, by falling obliquely on the fun, as mentioned above, must have forced off from his furface a quantity of matter equal to a 650th part of his body. This matter, being in a liquid state, would at first form a torrent, of which the largest and rarest parts would fly to the greatest distances; the smaller and more denfe, having received only an equal impulse, would remain nearer the fun; his power of attraction would operate upon all the parts detached from his body, and make them circulate round him; and, at the same time, the mutual attraction of the particles of matter would cause all the detached parts to assume the form of globes, at different diffances from the fun, the nearer moving with greater rapidity in their orhits than the more remote,

But it may be objected, that, if the plauets had been detached from the fun, they must have been burning and luminous, not cold and opaque bodies; nothing can have lefs refemblance to a globe of fire than a globe composed of earth and water; and, by comparison, the matter of the earth is totally different from that of the fun.

It may be replied to this objection, that the matter changed its form after its feparation, and that the fire, or light, was extinguished by the projectile motion communicated by the stroke. Befides, may it not be fupposed, that the fun, or a burning flar, moving with a rapidity equal to that of the planets, would foon be extinguished: and that this may be the reason why all the luminous, or burning ftars, are fixed, and without motion; and why those called new stars, which have probably changed their flations, are frequently extinguished and disappear? To confirm this remark, comets, when in their perihelia, ought to be inflamed even to their centre; but they never become luminous ftars; they only emit a burning vapour, a confiderable portion of which they leave behind them in their course,

In a medium which has little refistance, I acknowledge, that fire may fublift, although the burning body be moved with great rapidity. It must likewise be acknowledged, that what I have faid applies only to those stars which difappear for ever, not to those that appear and difappear at stated intervals, without changing their fituations in the heavens. Of these Maupertuis, in his discourse on the figure of the stars, has given a most fatisfactory account. But those which have appeared, and then vanished for ever, must unquestionably have been extinguifhed either by the quickness of their motion, or fome other cause. There is not a fingle example of a luminous star revolving round another; and not one of the fixteen planets which revolve round the fun have any light in themfelves.

Farther, fire, in fmall maffes, cannot fubfift fo long as in large ones. The planets would burn a confiderable time after they iffued from the fun; but, at length, would extinguish for want of combustible matter. For the same reafon, the fun itself will be extinguished; but at a period as much beyond that which extinguished the planets, as the quantity of matter in the fun exceeds that of the planets. However this may be, the separation of the planets from the fun, by the shock of a comet, appears sufficient to account for their extinction.

The earth and planets, when they iffued from the fun, were totally composed of liquid fire; in which flate they would continue no longer than the violence of the heat that kept them in fusion. But this heat would gradually decay from the moment they left the fun. During their fluid flate, they necessarily assumed circular figures; and their diurnal motion would elevate their equators, and flatten their poles. I agree with M. Leibnitz *, that this figure corre-

[.] Vid. Act. Brud. Lipf. an. 1602.

fronds fo exactly with the laws of hydroftatics, that the earth and planets mult necellarily have been once in a flate of fluidity occasioned by fire; and, consequently, that the interior parts of the earth mult be composed of vitrified matter, of which fand, free-stone, granite, and perhaus clay, are fragments, or forize.

It is therefore extremely probable, that the planets were originally parts of the fun feparated by a ftroke which communicated to them a projectile motion; and that their different diffances proceeded folely from the difference of their denlities. To complete this theory, it only remains to account for the diurnal motion of the planets, and the origin of their fatellites; which, inflead of adding fresh difficulties, will tend greatly to confirm my hypothesis: For rotation, or what is called diurnal motion, entirely depends on the obliquity of the stroke; an oblique impulse on the surface of a body necesfarily gives it a rotatory motion. If the body which receives the impulse be homogeneous, the rotatory motion will always be equal and uniform ; but it will be unequal, if the body confift of heterogeneous parts, or of parts different in denfity. Hence we may conclude, that the matter of each planet is homogeneous, because the diurnal motion of each is uniformly performed in the same time; and this circumstance is an additional proof, that portions of different denfities were originally feparated from the fun.

But the obliquity of the ftroke might be for great as to throw off fmall quantities of matter from the principal planet, which would neceffarily move in the fame direction. These parts. by mutual attraction, would reunite, according to their denfities, at different distances from the planet, follow its course round the fun, and at the fame time revolve about the body of the planet, nearly in the plane of its orbit. It is eafy to perceive that the portions we mean are the fatellites: Thus the formation, polition, and motion, of the fatellites correspond, in the most perfect manner, with our theory : for they all move in the fame direction, and in concentric circles round their principal planets, and nearly in the plane of their orbits. All thefe common effects, depending on an impulsive force, must have proceeded from a common cause, which was a projectile force communicated to them by the fame oblique stroke. This account of the motion and formation of the fatellites will be ftrongly supported, if the other circumstances and phænomena attending them be duly weighed. Those planets which are furnished with fatellites move quickeft round their axes. The revolution of the earth is quicker than that of Mars, in the proportion nearly of 24 to 15: the earth has a fatellite, and Mars has none; Jupiter, whose diurnal motion is 500 or 600 times more rapid than that of the earth, has four fatellites; and it is extremely probable, that Saturn, who

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We may even conjecture, with fome probability, that the plane of the equator of Saturn's ring is nearly the fame with that of the planet ; for, supposing, according to the preceding theory, the obliquity of the impulie which put Saturn in motion to have been very great, his diurnal motion would at first be in proportion to the excess of the centrifugal force above that of gravity, and, of course, a considerable quantity of matter would be thrown off from his equatorial regions, and necessarily assume the figure of a ring, the plane of which would be nearly the fame with that of his own equator. This quantity of matter, detached from the equatorial regions of Saturn, must have flattened the equator of that planet; which is the reason why, notwithflanding the rapidity with which we have fupposed him to revolve round his axis, the diameters of Saturn are not fo unequal as those of Jupiter, which differ from each other more than an eleventh part.

Though this theory of the formation of the planets and their fatellites appears to be extremely probable; yet, as every man has his own fandard of effimating probabilities of this nature, and as this flandard varies according to the different capacities of combining analogies more or lefs remote, I pretend not to convince those who are unwilling to believe. I have offered these

ideas to the public, not only because I thought them rational, and calculated to unravel a fitbject upon which, however important, nothing has hitherto been written; but because the impulsive motion of the planes giver infe to numberless phenomena in the universe, which admit not of an explanation by gravity alone. To those who may be displosed to deny the possibility of my theory, I would propose the following queries:

1. Is it not natural to imagine, that a moving body has received its motion from the impulse of some other body?

2. When feveral bodies move in the fame direction, is it not exceedingly probable, that they received this direction from a fingle flroke, or, at leaft, from flrokes every way fimilar?

3. When feveral bodies in motion have not only the fame direction, but are placed in the fame plane, is it not more natural to think that they received this direction and position from one impulse than from many?

4. Is it not probable, that a body put in motion by impulle, fhould receive it in an oblique direction; and configuently that it fhould be forced to move round its axis with a rapidity proportioned to the obliquity of the fircke? If these queries be not unreasonable, the theory of which we have given a fleeth will no longer have the anoperance of abstraction.

OF THE FORMATION

Let us now proceed to a more interesting object; let us examine the figure of the earth upon which fo many inquiries and observations have been made. As it appears, from the equality of the earth's diurnal motion, and the uniformity in the inclination of its axis, that it is composed of homogeneous parts which mutually attract each other in proportion to their quantities of matter; if its impulfive motion had been communicated in a direction perpendicular to its furface, it would necessarily have affumed the figure of a perfect fphere: But, having been flruck obliquely, it moved round its axis at the inflant it received its figure; and, from the combination of the projectile force and that of attraction, there resulted a spheroid figure, more elevated at the equator than at the poles; because the centrifugal force, arising from the diurnal rotation of the earth, must diminish the action of gravity, or that power which makes all the parts tend to the centre. Thus the earth, being composed of homogeneous parts, and having been endowed with a rotatory motion, must necessarily have assumed a spheroidal figure, the two axes of which differ from each other by a 230th part. To show that this is the real figure of the earth, we need not have recourse to hypotheses; it is capable of the clearest demonstration. The laws of gravitation are well known: That bodies attract each other directly as their quantities of matter, and inverfely as the figures of their diffances, admits not of a doubt. It can as little be doubted, that the total action of any body is composed of all the particular actions of its parts.

The parts of bodies are all mutually attracked in the above proportion; and all these attractions, when the body has no rotation, necessary by produce a fishere, and a fishered, when the body is endowed with a rotatory motion. This fishered is more or lefs flattened at the poles in proportion to the quickness of the diurnal motion; and the earth, in consequence of the celerity of its rotation, and the mutual attraction of its parts, has assumed the figure of a fisheroid, or which the two axes are to one another as 230 to 230.

Thus the earth, at the time of its formation, from the original conflictution and homegeneity of its parts, and independent of every hypothesis derived from the direction of gravity, took the figure of a fpheroid; and, from the known laws of mechanics, its equatorial diameter was necediarily elevated about fix lengues and a half more than its roles.

I shall dwell the longer on this article, because there are some geometers, who, from a system of philosophy they have adopted, and from a supposed direction of gravity, still imagine that the figure of the earth depends upon theory. The first thing to be demonstrated is the mutual attraction of the parts of matter: and the fecond, the homogeneity of the terreftrial globe. When these two facts are clearly proved, there will be no occasion to have recourse to any theory derived from the direction of gravity; because the earth's figure, in this case, must necessarily be as Newton determined it; and all the other figures affigned to it, in confequence of vortexes, and other hypotheses,

can have no existence. It will not be doubted, even by the most incredulous, that the planets are retained in their orbits by the power of gravity. The fatellites of Saturn gravitate towards that planet; those of Jupiter towards Jupiter; the moon gravitates towards the earth; and Saturn, Jupiter, Mars, the Earth, Venus, and Mercury, gravitate towards the fun. In the fame manner, Saturn, Jupiter, and the Earth, gravitate towards their refrective fatellites, and the fun gravitates towards the whole planets. Gravitation is there-. fore a general law, by which the whole planetary fystem is mutually affected; for action cannot exift without re-action. This mutual attraction of the planets is the law which regulates all their motions; and its existence is demonstrated by its effects. When Saturn and Jupiter are in conjunction, their mutual attraction produces an irregularity in their motion round the fun. The earth and the moon, also, mutually attract each other; but the irregularities in the moon's motion proceed principally from the attraction of the fun; and hence the fun, the earth, and the moon, mutually act upon each other. Now, the diffances are equal, is proportioned to their quantities of matter; and the fame power of gravity, which makes heavy bodies fall to the earth, and which extends as far as the moon, is likewife in proportion to the quantity of matter: The total gravity of a planet, therefore, is composed of the gravity of all its parts: Hence all the parts of matter, whether in the earth or planets, mutually attract each other; and, of course, the rotation of the earth round its axis must necessarily have bestowed on it the figure of a fpheroid, the axes of which are as 220 to 230. But the direction of gravity must be perpendicular to the earth's furface; and, confequently, unless the general and mutual attraction of the parts of matter be denied, no hypothesis derived from dation. But this mutual attraction, as we have feen, is demonstrated by actual observation: and the experiments made by pendulums prove its univerfal extension. No hypothesis, therefore, founded on the direction of gravity, can be admitted, without contradicting both reason and experience.

Let us now examine whether the parts compoling the terrestrial globe be homogeneous. I acknowledge, that, if the globe be supposed to F A

confift of parts differing in denfity, the direction of gravity would be different from that we have affigned, and that the earth's figure would vary according to the different suppositions which might be made concerning the direction of gravity. But, why make suppositions of this kind? Why, for example, do we suppose the parts near the centre to be more denfe than those more diffant from it? Are not all the particles which compose the globe united by their mutual attraction ? Every particle, therefore, is a centre; and there is no reason to believe that the parts which furround the centre are denfer than those which furround any other point. Befides, if any confiderable part of the earth were more denfe than another, the axis of rotation would approach nearer that part, and create an inequality in the diurnal revolution of the globe: It would produce an inequality in the apparent motion of the fixed flars; they would appear to move more quickly or flowly in the zenith or horizon, according as we happened to be fituated on the heavy or light parts of the earth; and the axis of the globe, not passing through its centre of gravity, would make a perceptible change in its polition. But nothing of this kind ever takes place. On the contrary, the diurnal revolution of the earth is equal and uniform. At every point of the earth's furface, the ftars appear to move with the fame quickness; and, if there be any nutation in its axis, it is too inconfiderable

OF PLANETS. to attract observation. Hence it may be concluded, that all the parts of the globe are at leaft nearly homogeneous.

If the earth were hollow, the cruft of which, for example, exceeded not three leagues in thickness, it would give rife to the following phænomena. 1. The mountains would bear fo great a proportion to the total thickness of the crust, that vast irregularities in the earth's motion would be occasioned by the attraction of the moon and of the fun: When the moon was in the meridian of the more elevated parts, as the Cordeliers, her attraction upon the whole globe would be much greater than when the was in the meridian of the lower parts. 2. The comparative attraction of the mountains would be greatly increased; and the experiments made on Mount Chimboraco, in Peru, would have given more degrees in the deviation of the plumbline than they actually gave feconds. 3. The weight of bodies would be greater on the tops of mountains than in the plains; and men would find themselves more weighty, and would walk with more difficulty, in high than in low grounds. These observations, and many others which might be made, should convince us, that the interior parts of the earth are not hollow, but that they are composed of matter of a considerable denfity.

If, on the other hand, the earth, at the depth of two or three leagues, confifted of matter much

much denfer than that we are acquainted with, upon descending even into ordinary pits, we fhould find ourselves considerably heavier; and the motion of pendulums would there be more accelerated than they actually are when brought down from a hill to a plain. Hence we may prefume that the interior parts of the earth confift of matter nearly fimilar to that on its furif we confider that the earth, at the time of its original formation, when it assumed its present foheroidal figure, was in a flate of fusion, and, confequently, that all its parts were homogeneous, and nearly of equal dentity. The matter on of the interior parts, has, in the revolutions of time, undergone many changes from external causes; and to these are to be ascribed the proties. But it ought to be remarked, that the denfest bodies, as gold, and other metals, are most rarely to be met with; and, consequently, that the greatest quantity of materials at the furface, have fuffered little alteration with regard to denfity. The most common materials, indeed, as fand and clay, differ fo little in denfity, that we may conjecture, with much probability, the internal parts of the earth to confift of a vitrified matter, the denfity of which is nearly equal to that of fand; and, confequently,

that the whole globe may be confidered as one

But, it may be faid, that, though the earth were composed of encentric best, of different denfities, the equality of its diurnal motion, and the uniform inclination of its axis, would remain equally undfuthed, as upon the disposition of its confuting wholly of homegeneous matter. This fallow; but I demand, at the fame time, whether there be any reason for believing that these best of different dendities really exist? Whether this method of foliving difficulties be not an attempt to adjust the works of mature to our own imaginations? And whether suppositions, eather founded on obferration nor analogy, ought to find admittance into exherted were cause?

Hence, it is apparent, that the earth received its fisheroidal figure in enfequence of its diurnal motion, and the mutual attraction of its parts; that this figure necessarily related from the globe's being in a liquid flate; that, according to the laws of gravity and of a centrifugal force, it could not polifibly assume any other figure; that, at the moment of its formation, the difference between its two diameters was, as at prefent, equal to a 250th part; and, of course, that all other hypothetes which make this difference greater or left, are mere fictions, and delered

Perhans

Perhaps it may be objected, that, if this theory be well founded, and if the proportion of the axes of the two diameters be as 229 to 220, how came the mathematicians fent to Lapland and Peru to concur in making it as 174 to 173.7 Why fhould finch a difference fubfilt between practice and theory? And, is it not more reasonable to give the preference to actual measurement, especially when executed by the ableft mathematicians in Europe*, and furnished with all the necessary armanatus?

To this I reply, that I mean not to combat the observations made at the equator, and near the pole: that I doubt not of their exactness: and that the earth may actually be elevated a 175th part more at the equator than at the poles. Still, however, I maintain my theory; and I perceive clearly how it may be reconciled to practice. The difference between the two conclusions is about four leagues in the two axes. The equatorial regions are found to have an elevation of two leagues more than they ought to have by the theory. This height of two leagues corresponds exactly with the greatest inequalities which have been produced on the furface of the globe by the motion of the fea, and the action of fluids. Here fome illustration is necessary. At the time of the earth's formation, in confequence of the mutual attraction of its parts, and of its centrifugal force, it must have assumed a

M. de Maspertuis, figure de la terre.
 fpheroidal

fpheroidal figure, with its axes different by a 230th part. This would be the real figure of the earth while it remained in a state of liquefaction. But, after cooling for fome time, the rarified vapours, like those in the tail or atmofphere of a comet, would condense, and fall on the furface in the form of air and water; and, when these waters begin to be agitated by a flux and reflux, fand, and other bodies, would be gradually transported from the poles towards the equatorial parts. This operation, when continued for fome time, would necessarily fink the poles, and elevate the equator in the fame proportion. The furface of the earth being likewife exposed to the winds, to the action of the air and of the fun ; all thefe caufes would concur with the tides in furrowing the earth, in scooping out valleys, in elevating the mountains, and in producing other fuperficial irregularities, none of which, perhaps, exceed a league in thickness, even at the equator. This inequality of two leagues may be supposed to be the greatest that can take place on the surface; for the highest mountains exceed not a league in height; and the most profound parts of the ocean, it is probable, are not above a league in depth. Thus my theory perfectly coincides with practice. The earth's equator could not, at first, be elevated more than fix leagues and a half above the poles; but the changes produced on the furface might give it a ftill greater elevation. Natural hiftory

OF THE FORMATION

history wonderfully fupports this opinion; for, in the preceding diffeourie, we have proved, that, from the tides and other motions of the waters, have proceeded mountains, and all the other inequalities on the furface of the globe; and that, at great depths, as well as upon the greatest heights, bones, fhells, and other releas of fea and had animals, have been diffeovered.

From what has been observed, it may be conjectured, that, in order to find primitive earth, and substances which have never been removed from their original stations, we must dig in countries near the poles, where the hed of unmoved earth will be thinner than in fouthern

In fine, if the measurement by which the figure of the earth was determined be firifly ferutinized, we shall find that it is not altogether free from hypothetical reasoning: For it proceeded on the supposition that the earth was a regular curve: But, as the earth is liable to confiderable and constant changes from a thousand causes, it is impossible that it could have retained any perfectly regular figure; and hence, agreeable to our theory, and the opinion of Newton, the poles might originally be only flattened a 230th part. Befides, though we have the exact length of a degree at the equator and polar circle, yet we have not the exact length of a degree in France: and the measures of M. Picard have never been confirmed. It may be

added,

added, that the diminution and increase in the motion of the pendulum agree not with the conclusions drawn from measurement; and that, on the contrary, they correspond very nearly with the theory of Newton. These circumstances tend farther to convince us, that the poles are not depreffed above a 230th part; and that, if there be any difference, it can proceed from nothing but the inequalities produced on the furface by the waters, and other external causes. But these inequalities are by no means fo regular as to justify any hypothesis, which supposes the meridians to be ellipses, or any other perfect curves. Hence it appears, that, though many degrees should be successively measured in different regions, we cannot, by that alone, afcertain the exact depression of the poles, nor determine how much it exceeds or falls thort of a 2 20th part.

May we not likewife conjecture, that, if the inclination of the earth's axis has been changed, this effect could not be produced but by the changes on the furface, fince all the other parts are homogeneous; that this variation is, of courfe, too finall to be perceived by affronomers; and that, if the earth be not diffurbed by a comet, or flone other extremal caufe, its axis will for ever preferve its prefent and original inclination?

Not to omit any conjecture that feems reafonable, may we not suppose, that, as the mountains and other inequalities on the surface of the earth, have originated from the action of the tides, those which we perceive in the moon have been produced by a fimilar cause? The mountains of the moon are indeed higher than those of the earth; but her tides are likewise ftronger; because the earth, the fize of which is much larger, raifes the tides of the moon with a fuperior force. This effect would be greatly augmented, if the moon, like the earth, had a quick diurnal motion. But, as the moon uniformly presents the same face to the earth, the tides are raifed only in proportion to the motion occasioned by her librations, which alternately expose to our view a fmall fegment of her other hemisphere. This cause, however, must produce tides very different from those of our feas: and their effects will, of courfe, be much less confiderable, than if the moon had possessed a diurnal revolution round her axis, equally quick as the rotation of the earth.

I thould compofe a volume equal to that of Barnet or Whilton, were I to extend the ideas prefented by the above theory; and were I, in imitation of the laft-mentioned author, to clothe them in a geometrical drefs, I might add confiderably to their importance. But I have always thought, that hypothefs, however probable, deferve not to be treated fo pompoully. It is apt to give them the air of quackery and imposition.

PROOFS

OF THE

THEORY OF THE EARTH.

ARTICLE II.

Of the System of Whiston *.

This author begins his theory with a differentian on the creation of the world. He alleges, that the account given of it by Mofes is not properly underflood; and that, in inquiries of this kind, men, contenting them-felves with the molt evident and siperficial views, give too little of their attention to nature, restron, and philosophy. The common notions, he observes, concerning the fix days work, are false; and the description of Mofes is not an exact or philosophic account of the creation and origin of the universe, but only an hildroisi

* See a new Theory of the Earth by Will. Whifton, London, 1708.

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